

Scintillation-based background rejection methods

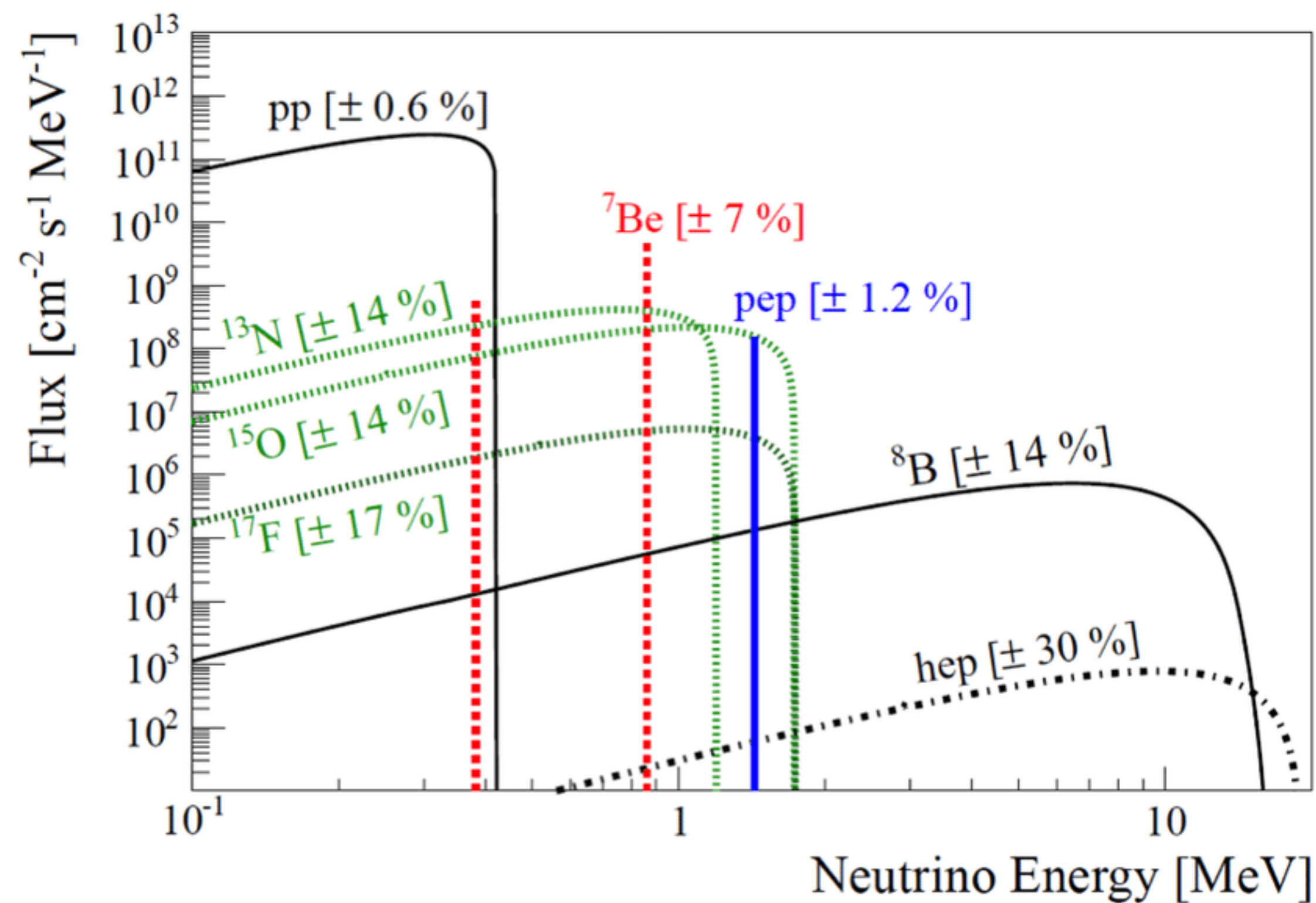
Anyssa Navrer-Agasson

DUNE Low Energy Physics Working Group - August 11, 2021

Solar neutrinos in DUNE

Solar neutrinos

DUNE potential



DUNE could deliver world-leading results in solar neutrinos:

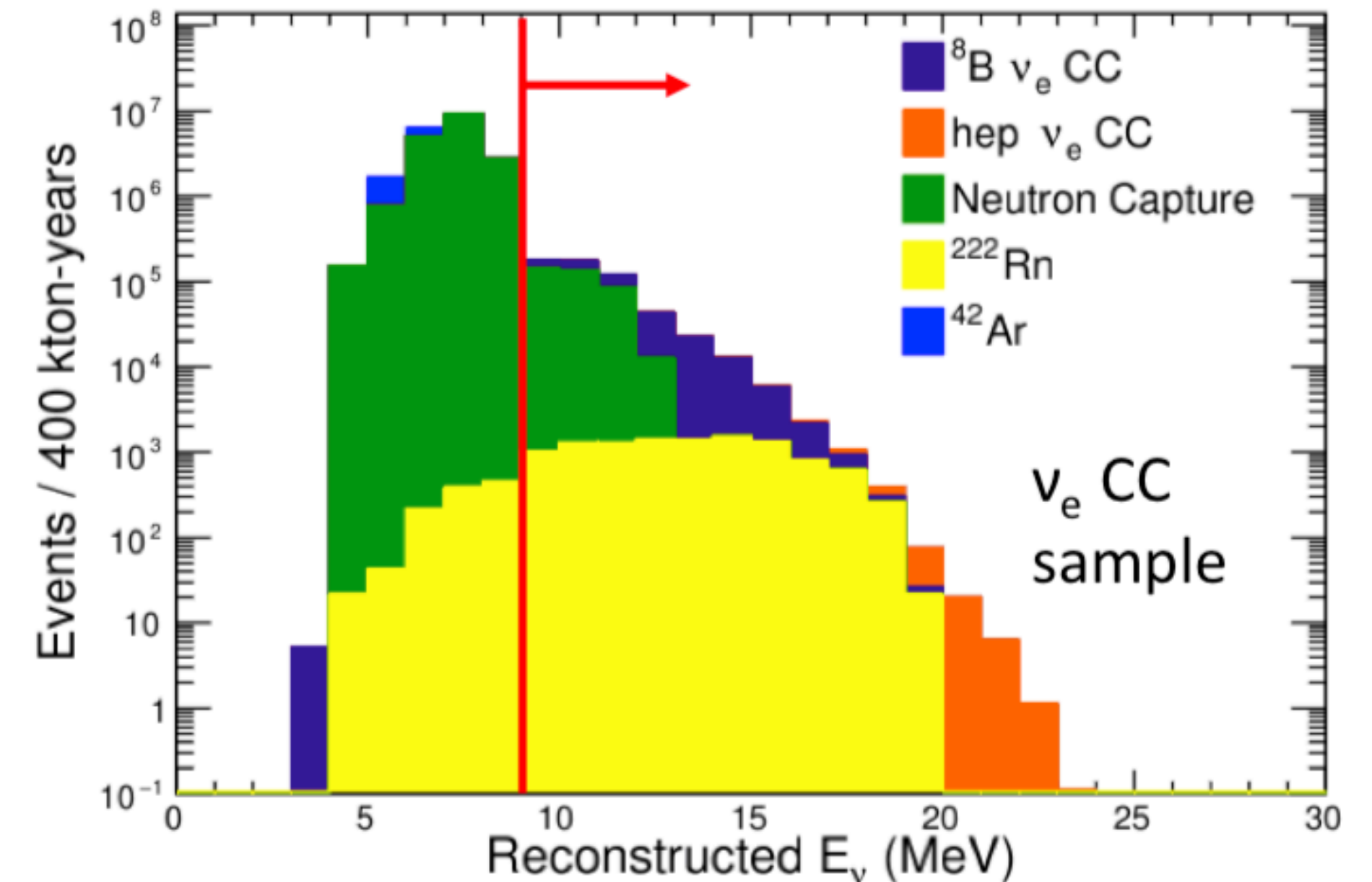
- Precision measurement of the ^8B flux
- First observation of the hep flux
- Neutrino-mixing parameters measurements

Solar neutrinos

Previous study

- Dan Pershey studied charge signal from solar neutrinos and background sources
- Identified $^{40}\text{Ar}(\alpha, \gamma)$ reactions from ^{222}Rn decays as a concerning background
 - ^{222}Rn uniformly distributed in the LAr
 - α from Rn decay chain absorbed by ^{40}Ar with subsequent γ emission (rare, but too frequent to be negligible)
 - $E_\gamma \sim 15 \text{ MeV}$

Assume 10 mBq/kg of ^{222}Rn
(see Juergen's talk)



Idea: use scintillation light to reduce ^{222}Rn backgrounds

Solar neutrinos

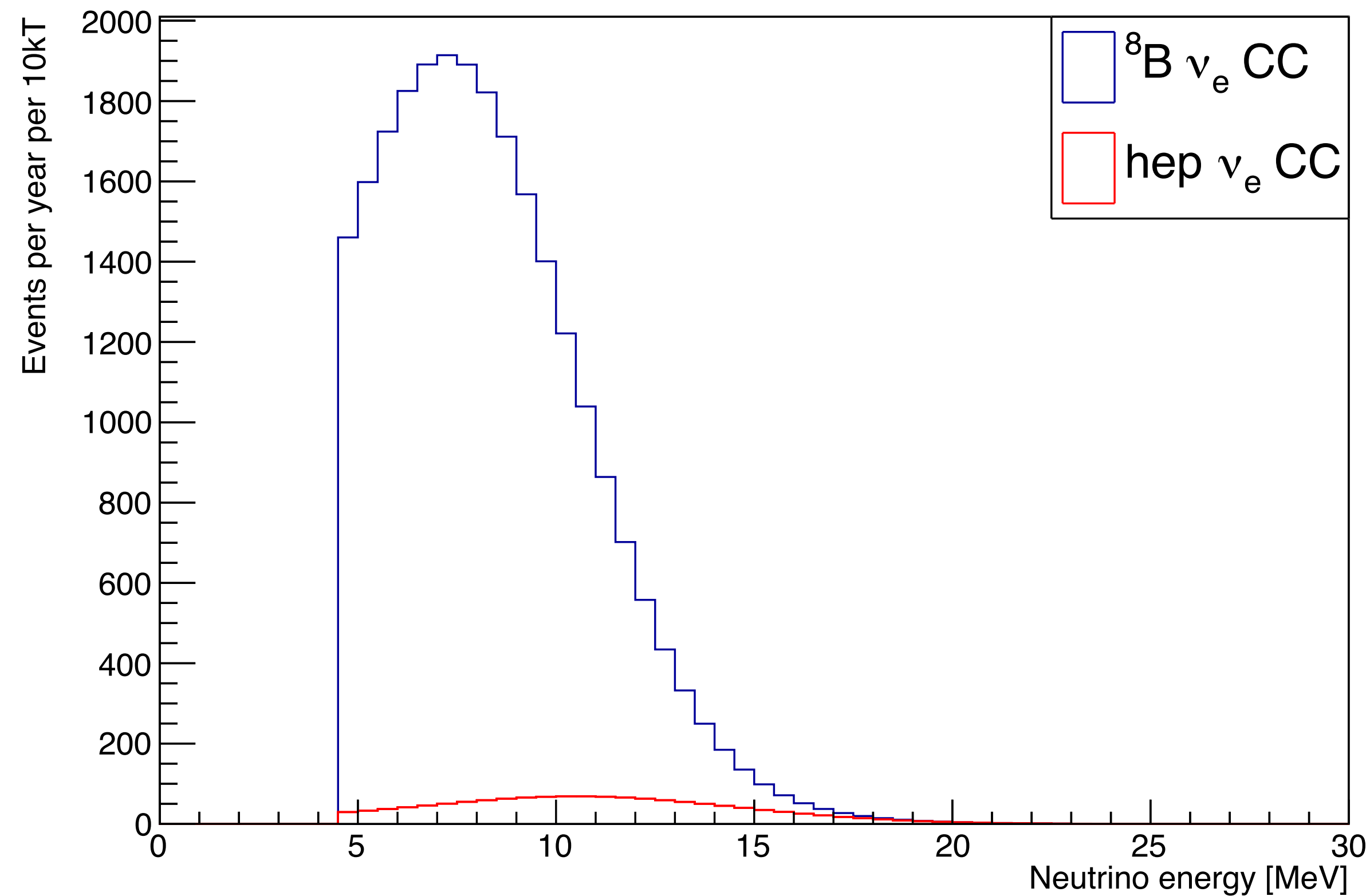
Our study

- Improve solar neutrino signal selection with scintillation-based methods
 - Make use of Pulse Shape Discrimination (widely used in dark matter LAr experiments)
 - Rely on non-uniform distribution of ions from ^{222}Rn decays in the TPC
 - No charge information

Detector and simulation

Detector and simulation

Event rate in DUNE



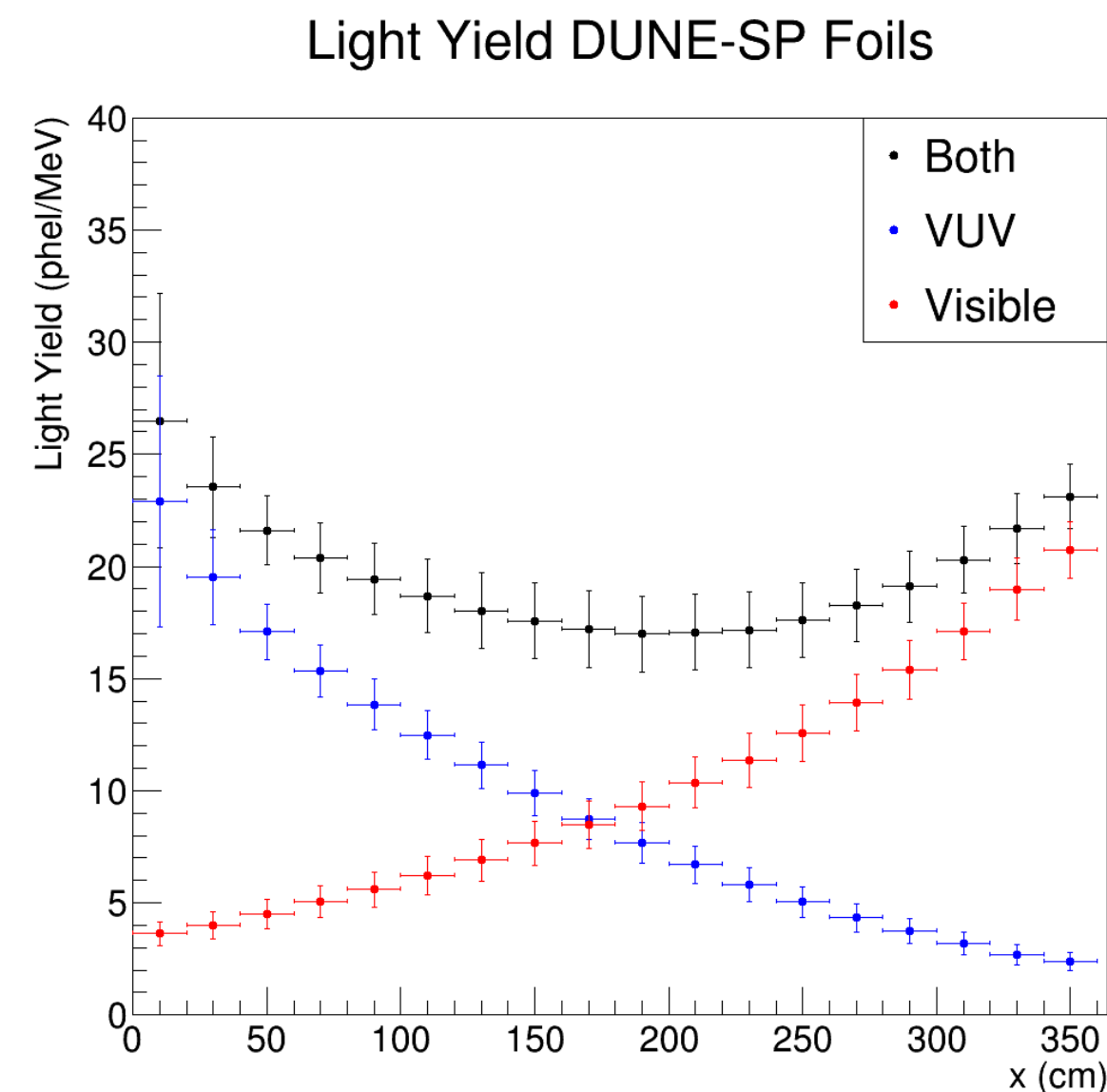
Initial fluxes from
E.Vitagliano, I.Tamborra, G. Raffelt

Event rates are computed using SNOwGLoBES (see [website](#))

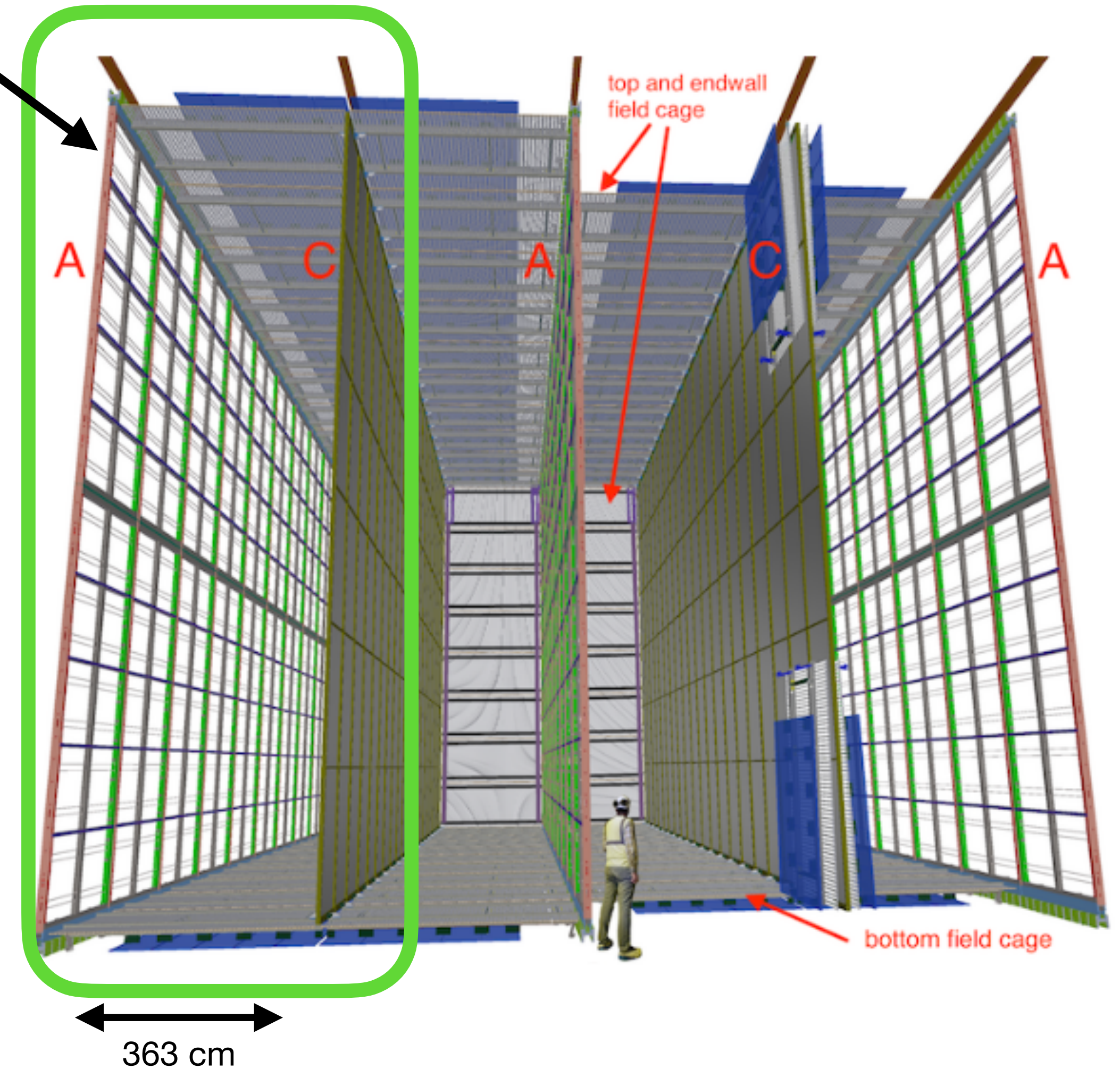
Detector and simulation

Geometry

- Simulate events in one section of module, between anode and cathode
- Study impact of Xe doping
- Add TPB coated foils at the cathode plane to enhance light collection



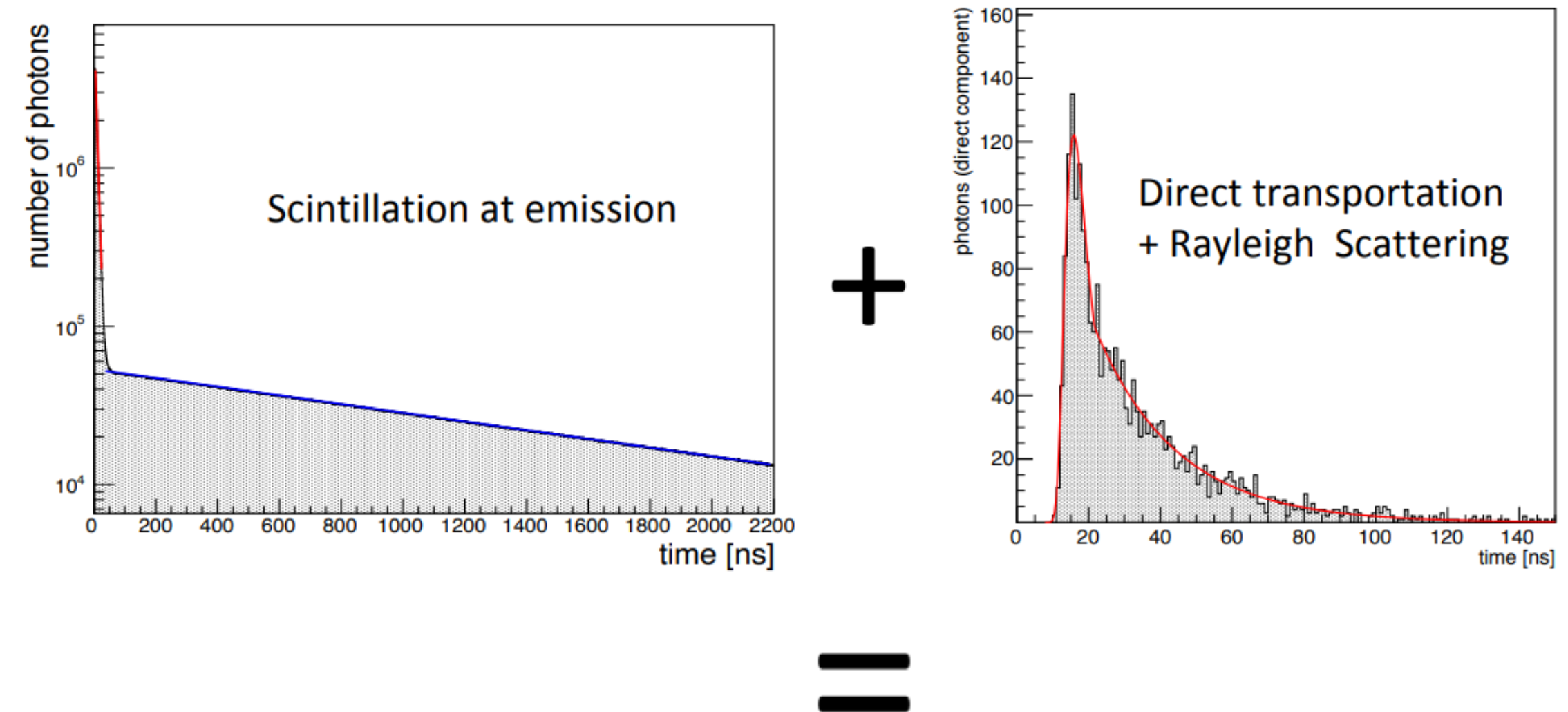
Photon detectors



Detector and simulation

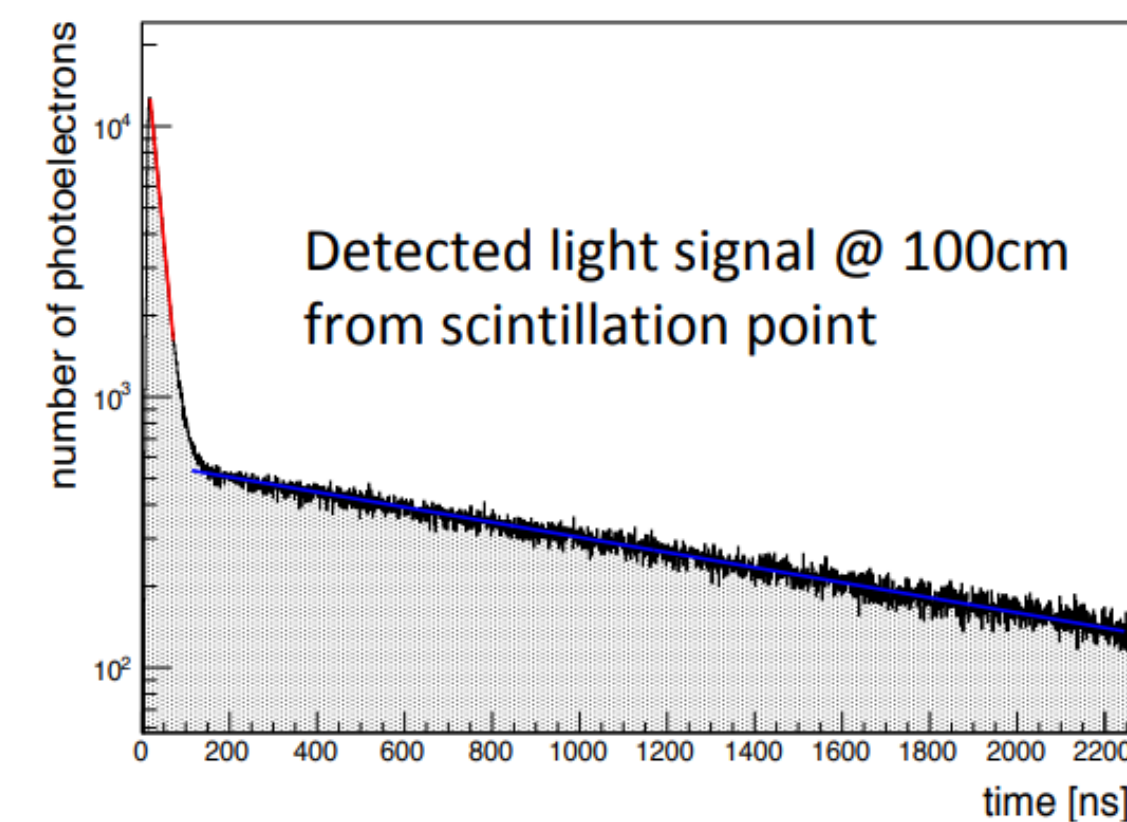
Semi-analytic scintillation model

- Geometric prediction of number of incident photons and direct propagation time
- Parametrised corrections accounting for effects of Rayleigh scattering, absorption and reflections



Standard fast optical simulation in DUNE-SP, SBND

(see [arXiv:2010.00324](https://arxiv.org/abs/2010.00324) for further details)

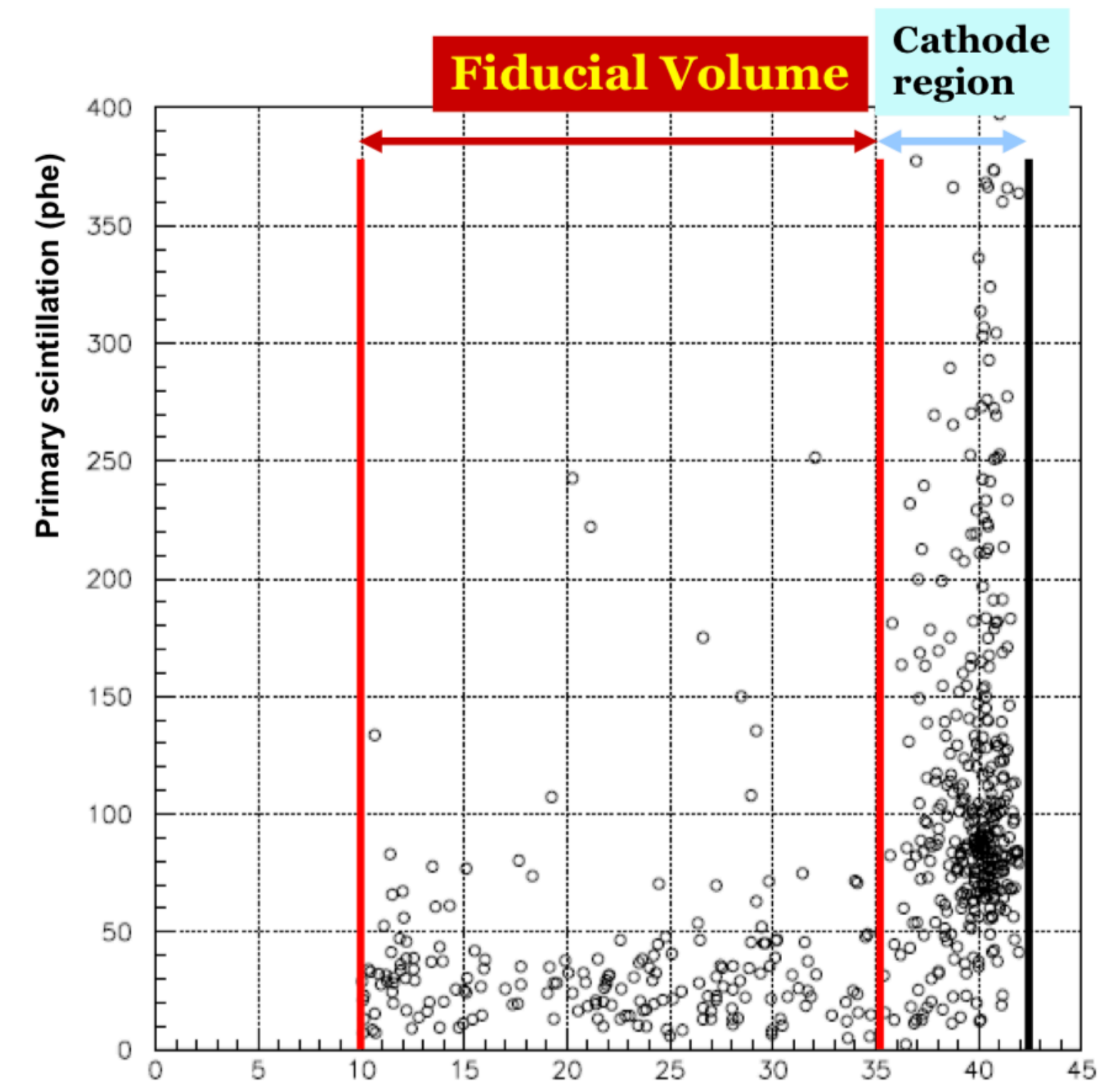


Ion migration model

Ion migration model

Context

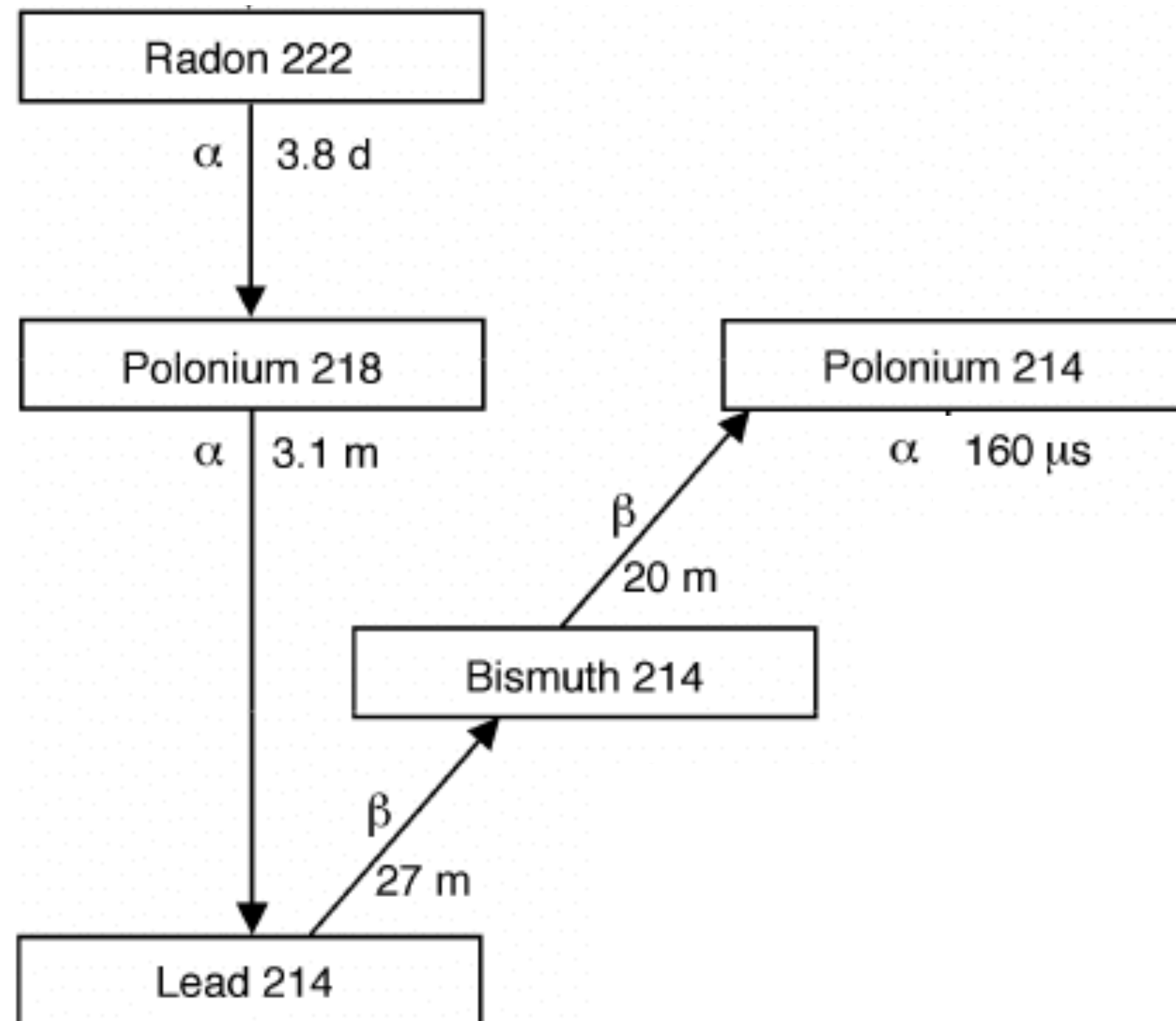
- Some dual phase TPC experiments observed an accumulation of ^{222}Rn and daughters events near the cathode
 - Due to some radon daughters being produced as positive ions
- Can provide a way to discriminate against background if we are able tag events near the cathode
- Need to know the fraction of alpha decays happening at the cathode



Observation by WArP

Ion migration model

Method



- Simulate the decay position of the different elements of the chain
 - The two key parameters are the ionisation fraction and drift velocity
- ^{222}Rn distributed uniformly along x-position with no drift
- ^{214}Po drift neglected due to its short half-life
- Extract the x distribution of alpha decays

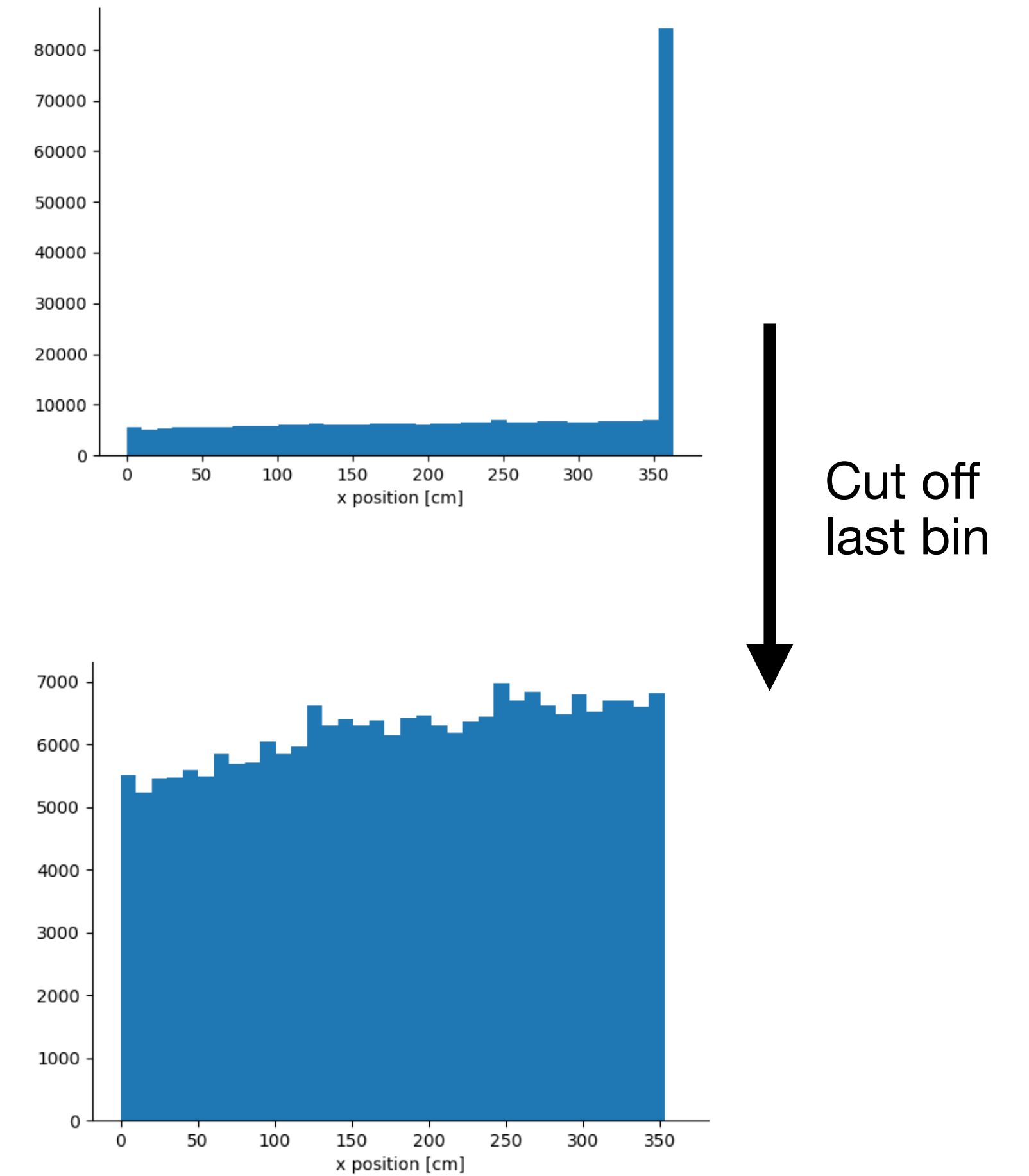
Ion migration model

Decay positions

Isotope	Drift velocity [cm/s]	Ion fraction
Po218	0.43	0.37
Pb214	0.4	0.37
Bi214	0.4	0.56

Distance to the cathode	Fraction of alpha decays [%]
< 30cm	34.72
< 20cm	32.5
< 10cm	30.31

Distribution used as input to simulate $^{40}\text{Ar}(\alpha,\gamma)$ decay positions



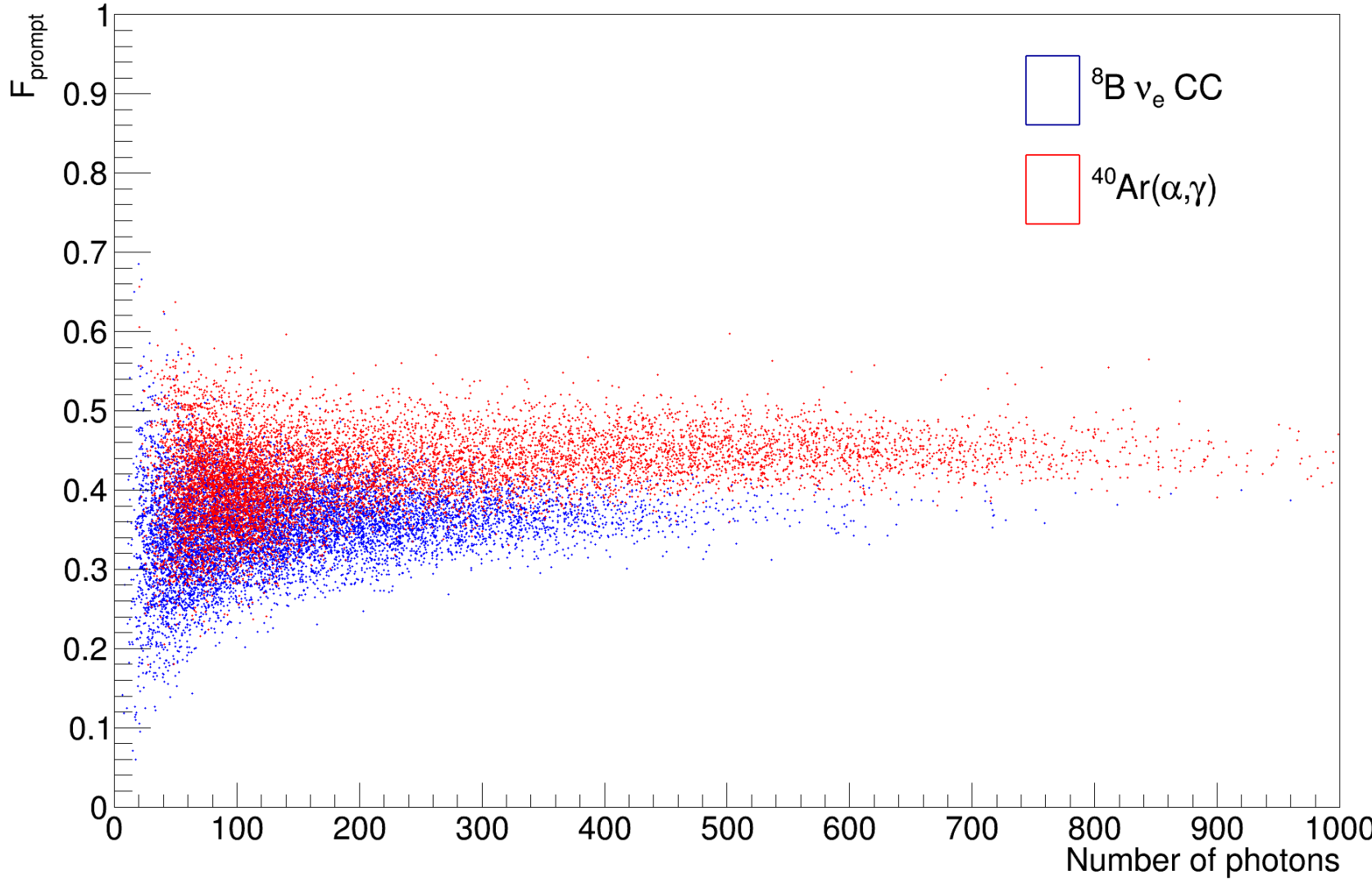
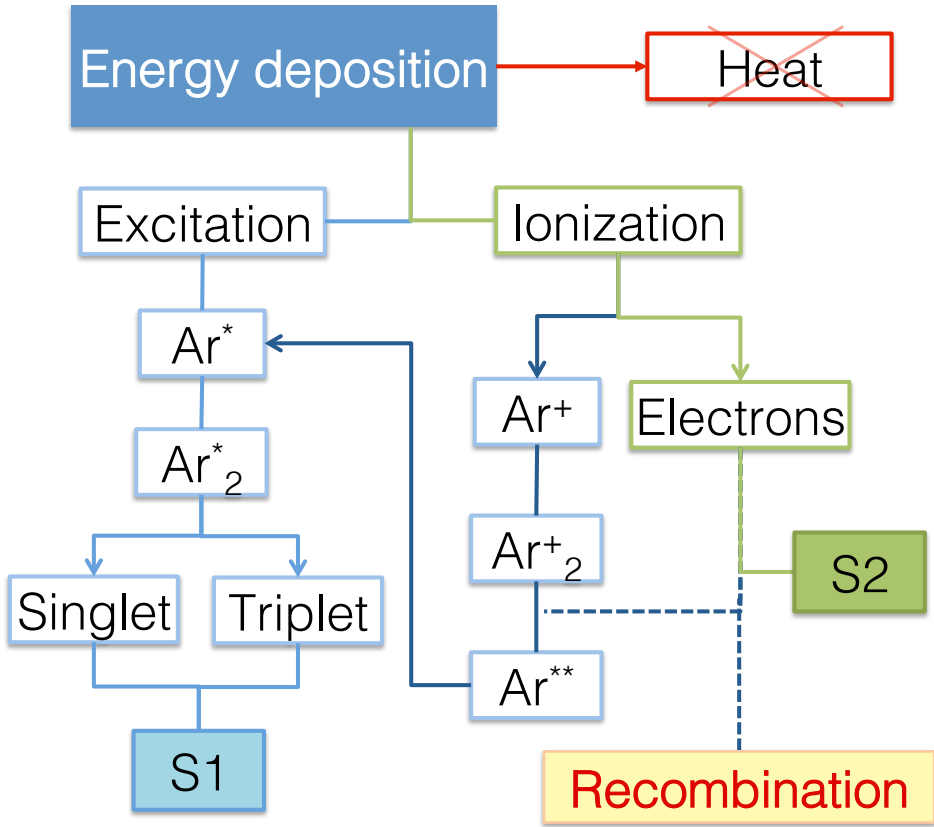
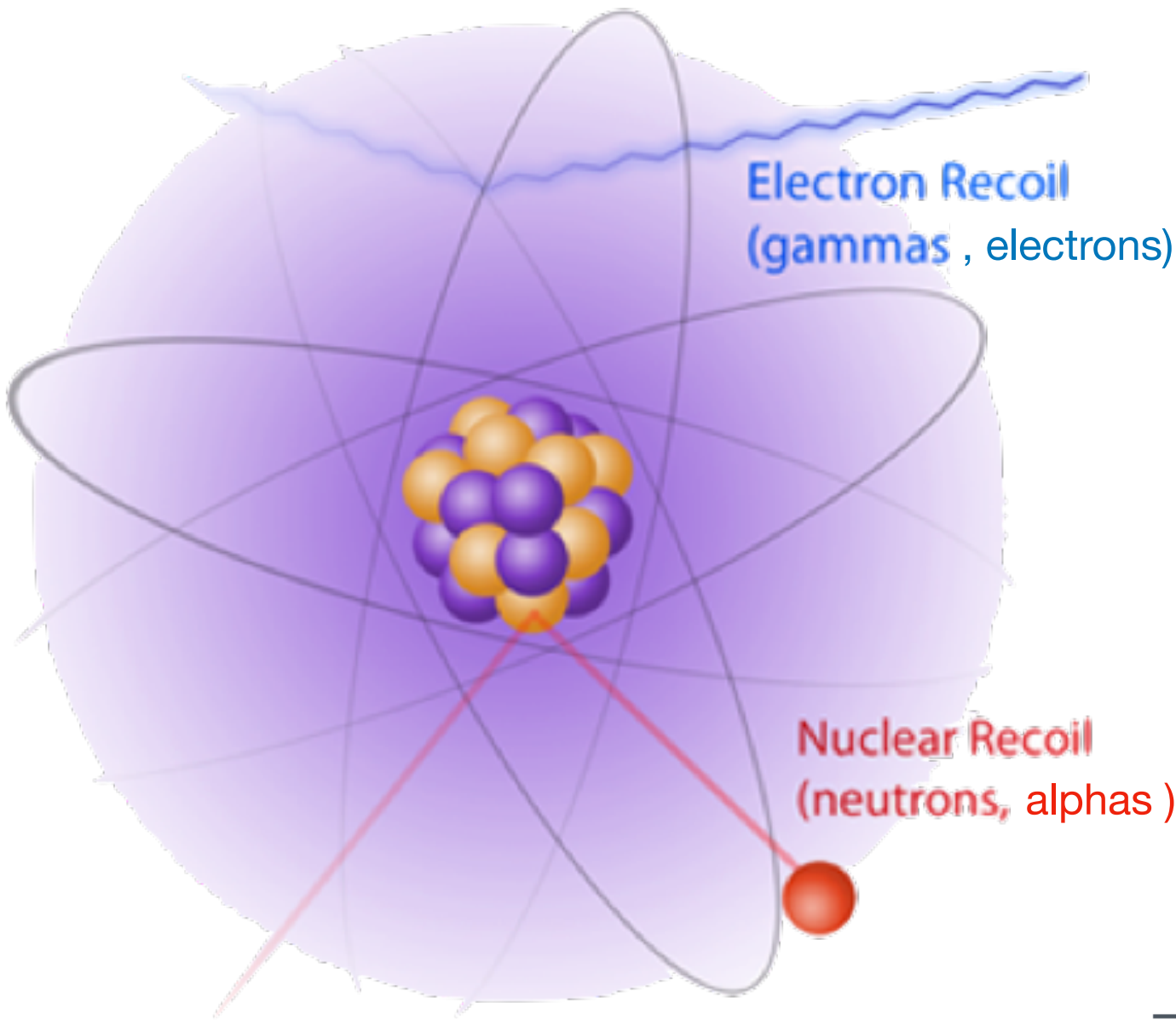
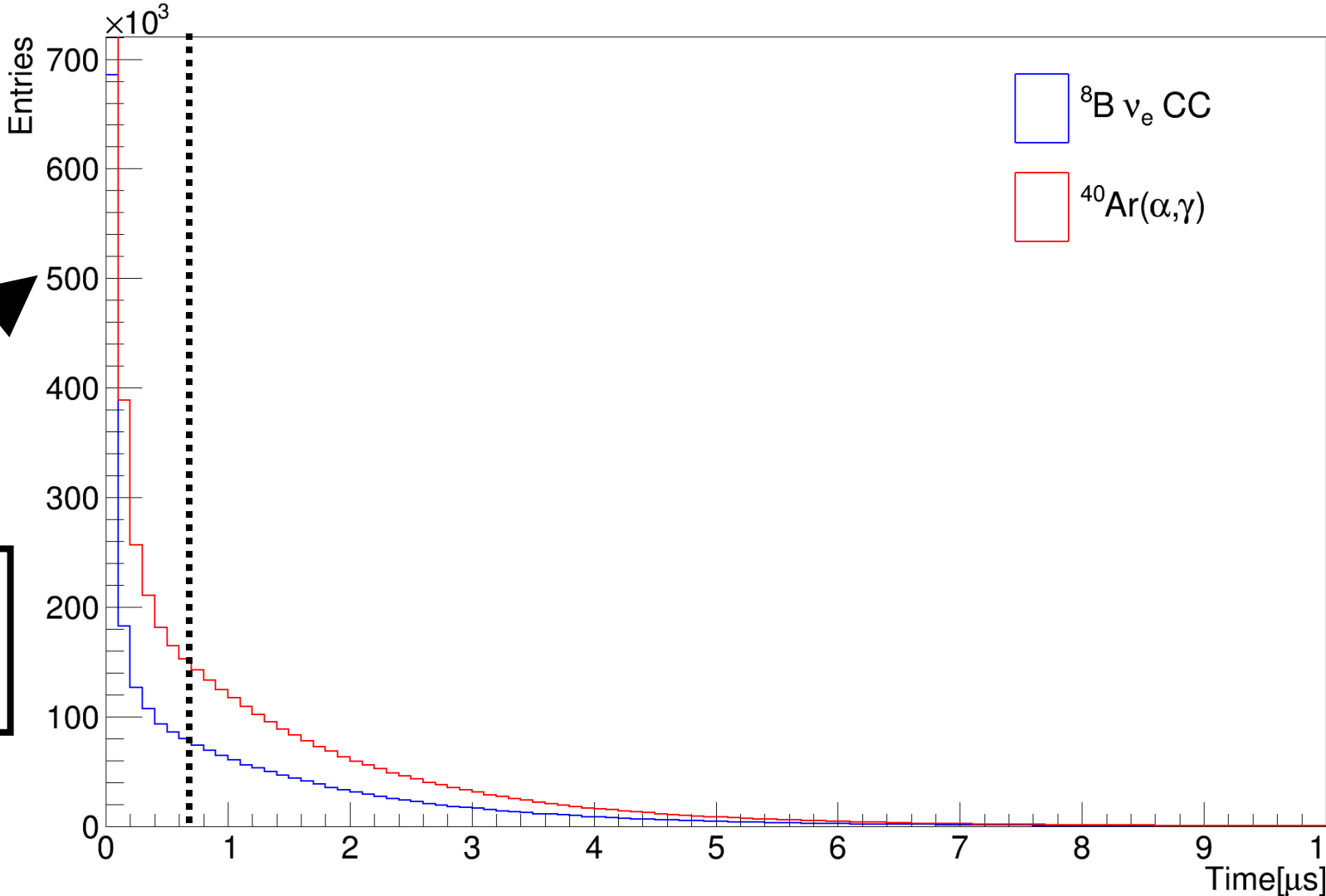
For more details on this study, see [full collaboration meeting talk](#)

Pulse shape discrimination

Pulse shape discrimination

Principle

PSD parameter f_{prompt} :
fraction of light seen in a certain time window

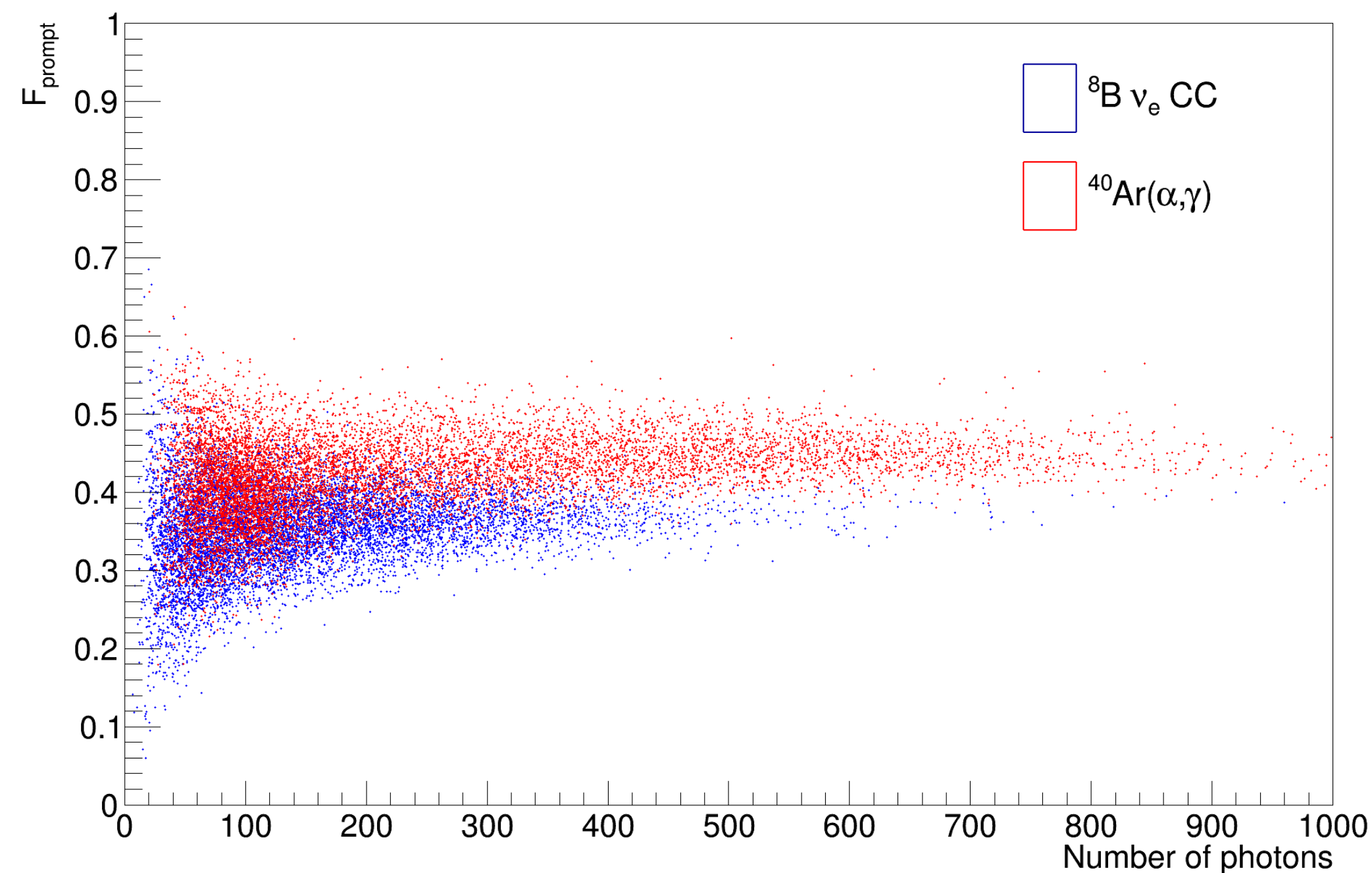


	Singlet	Triplet
Time constant	$\sim 7 \text{ ns}$	$\sim 1.6 \text{ } \mu\text{s}$
Population ratio for Electron ionizing	33%	67%
Population ratio for Nucleus ionizing	75%	25%

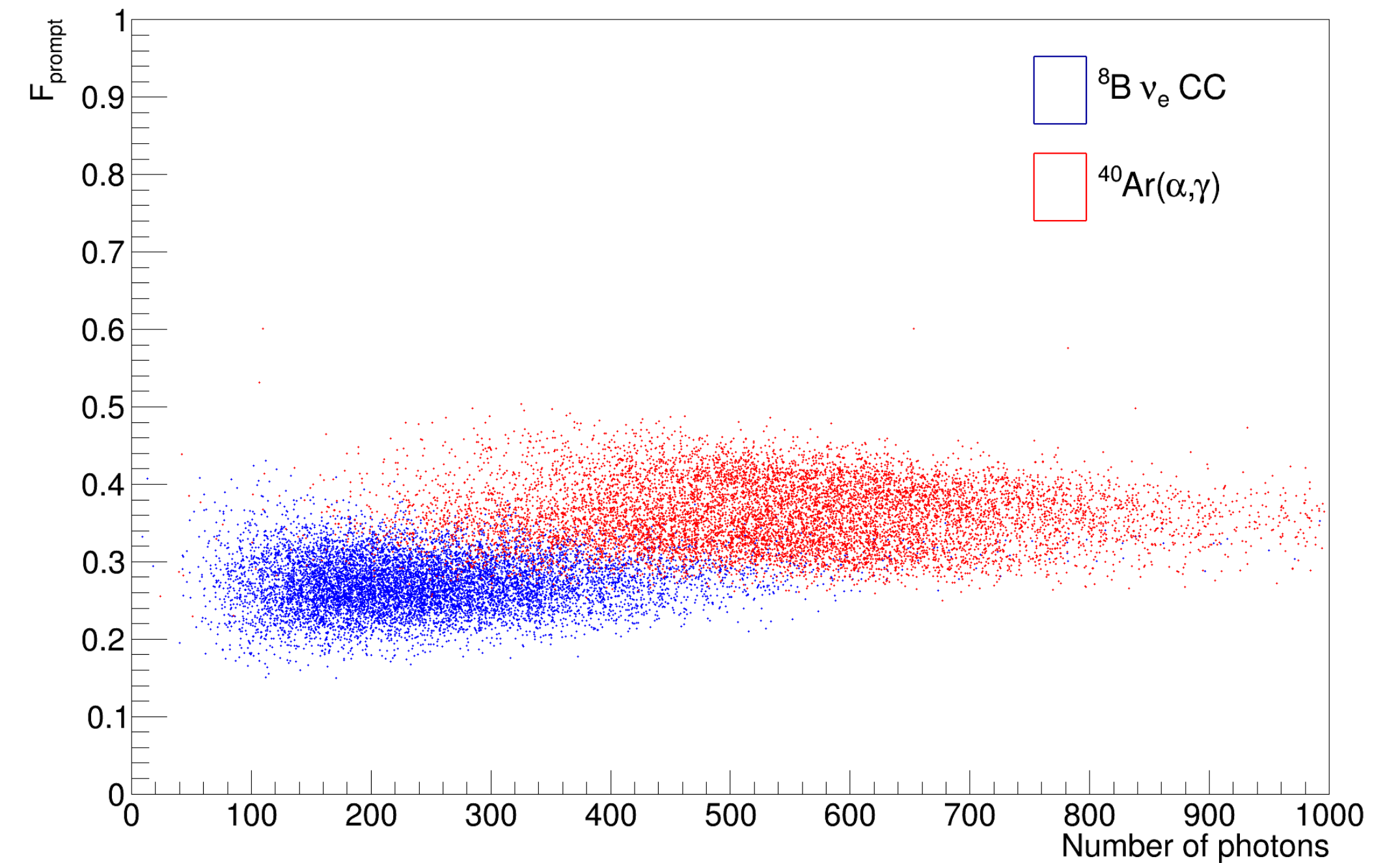
Pulse shape discrimination

Adding reflective foils

Without foils



With foils



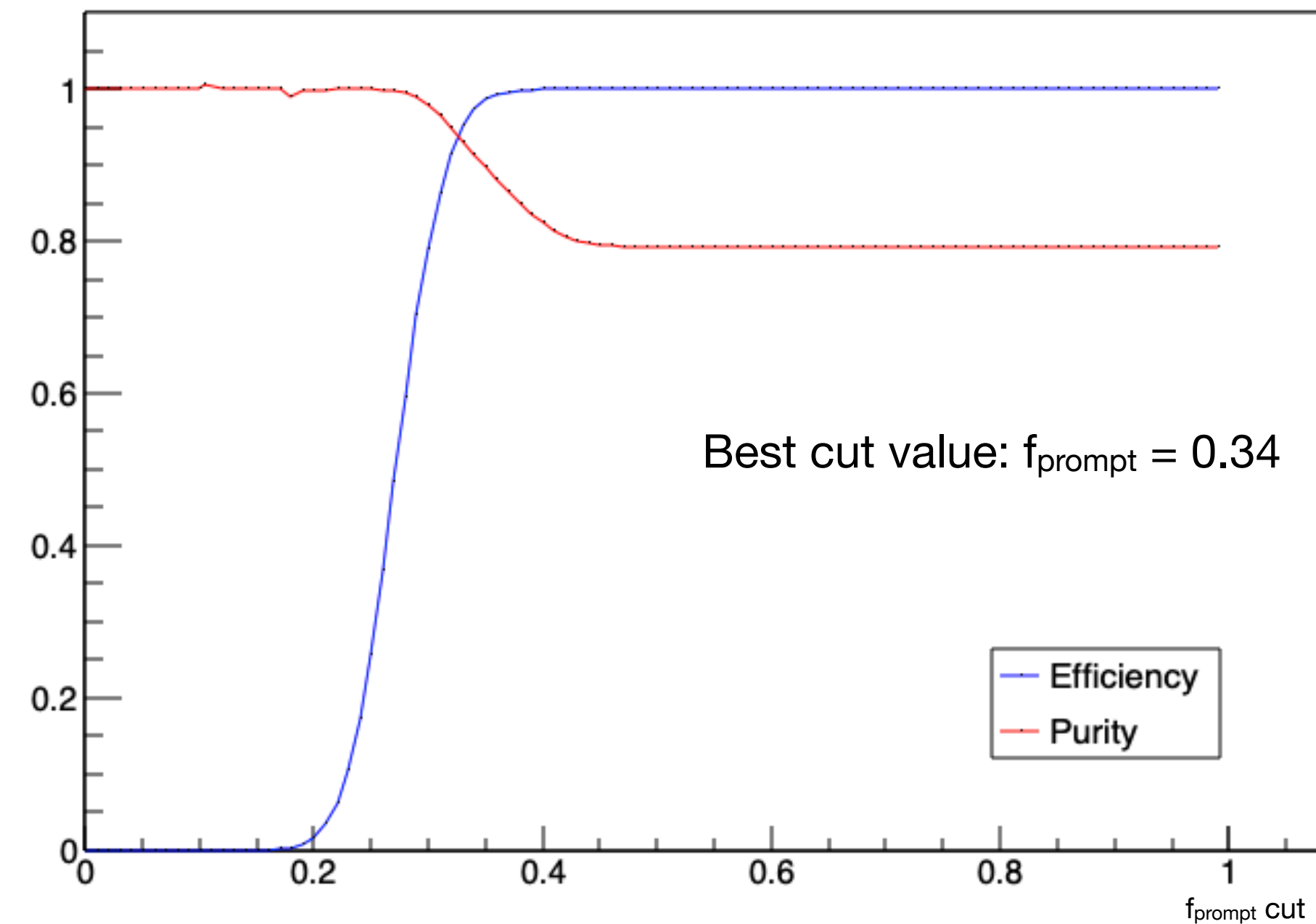
Foils improve separation as a large fraction of the ${}^{40}\text{Ar}(\alpha, \gamma)$ events happen near the cathode

- Not enough photons collected for efficient PSD without boost from foils

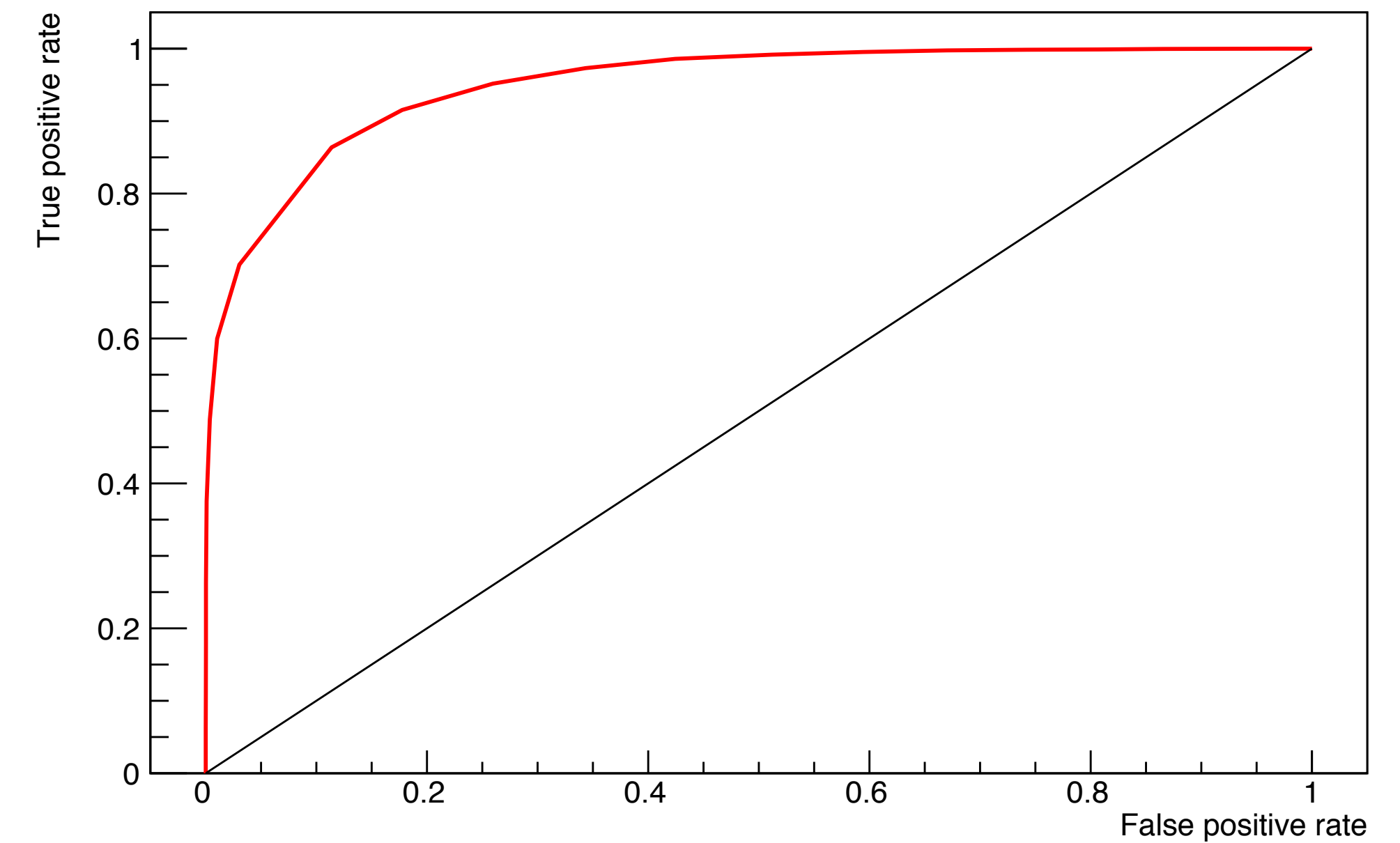
(all subsequent plots will be in the configuration with foils)

Pulse shape discrimination

F_{prompt} cut performance



Efficiency and purity



ROC curve

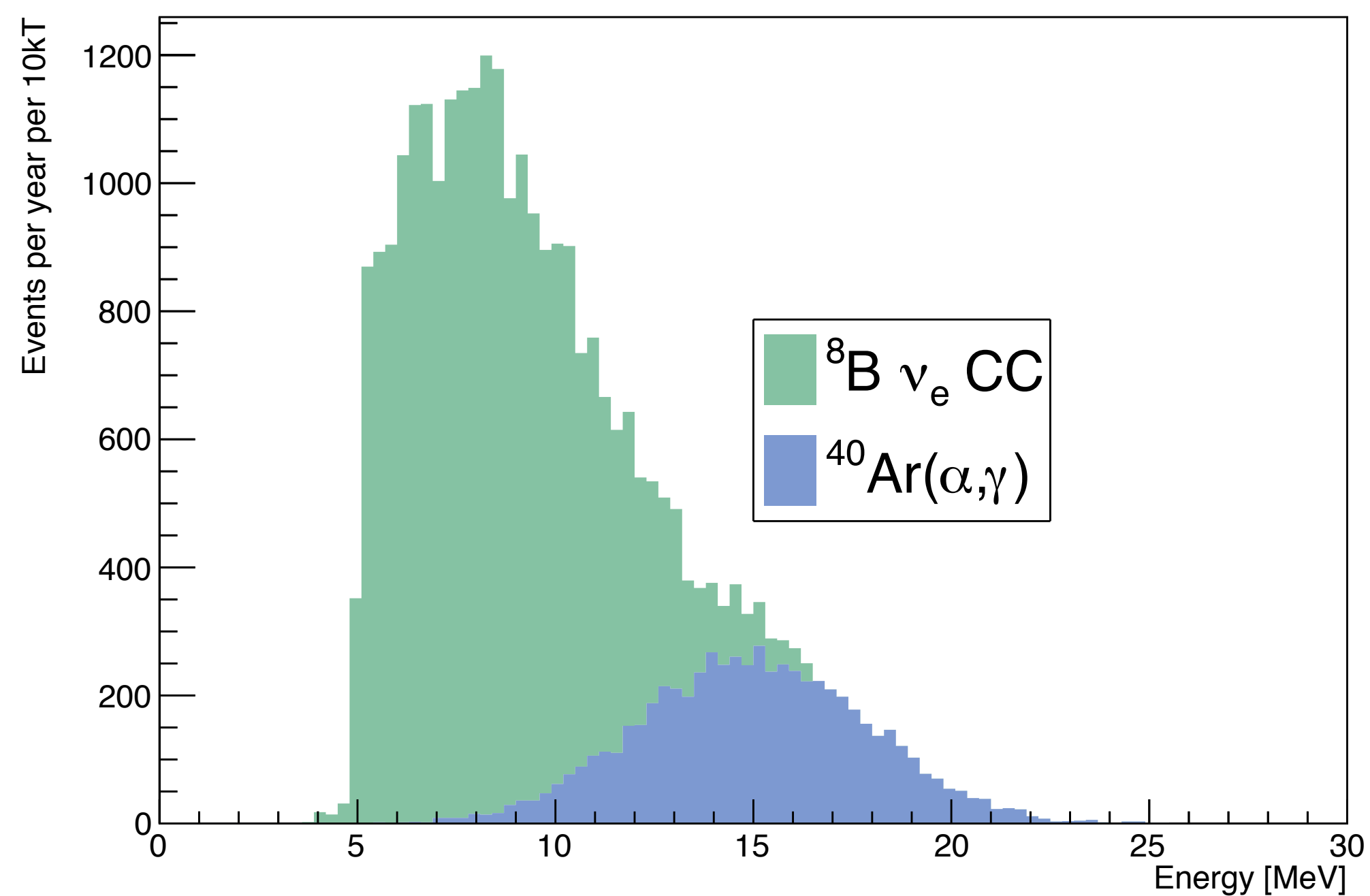
Pulse shape discrimination

Selected events

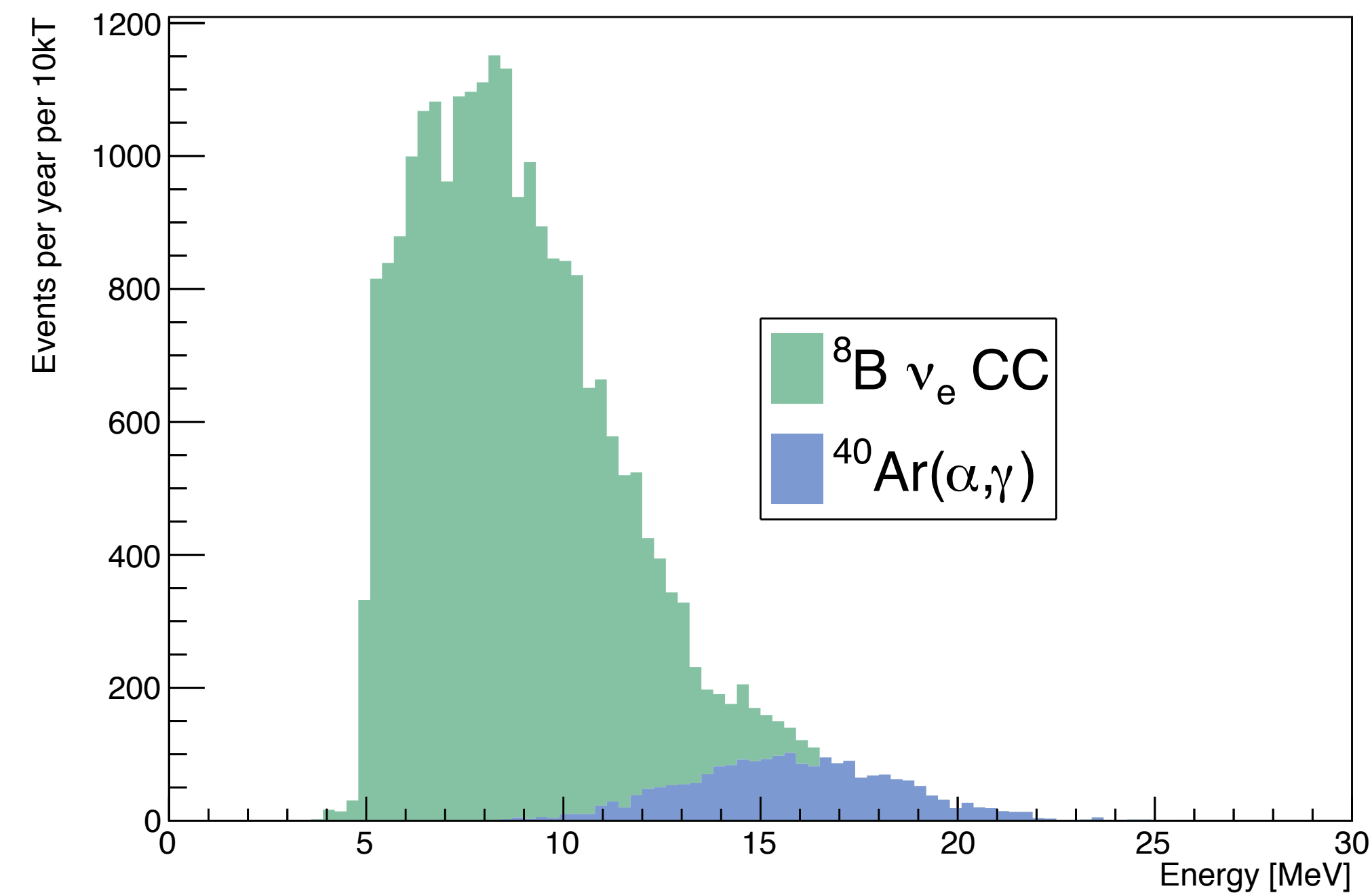
Signal efficiency: 97.3 %

Signal purity: 91.5%

Background efficiency: 34.5%



Before f_{prompt} cut



After f_{prompt} cut

Pulse shape discrimination

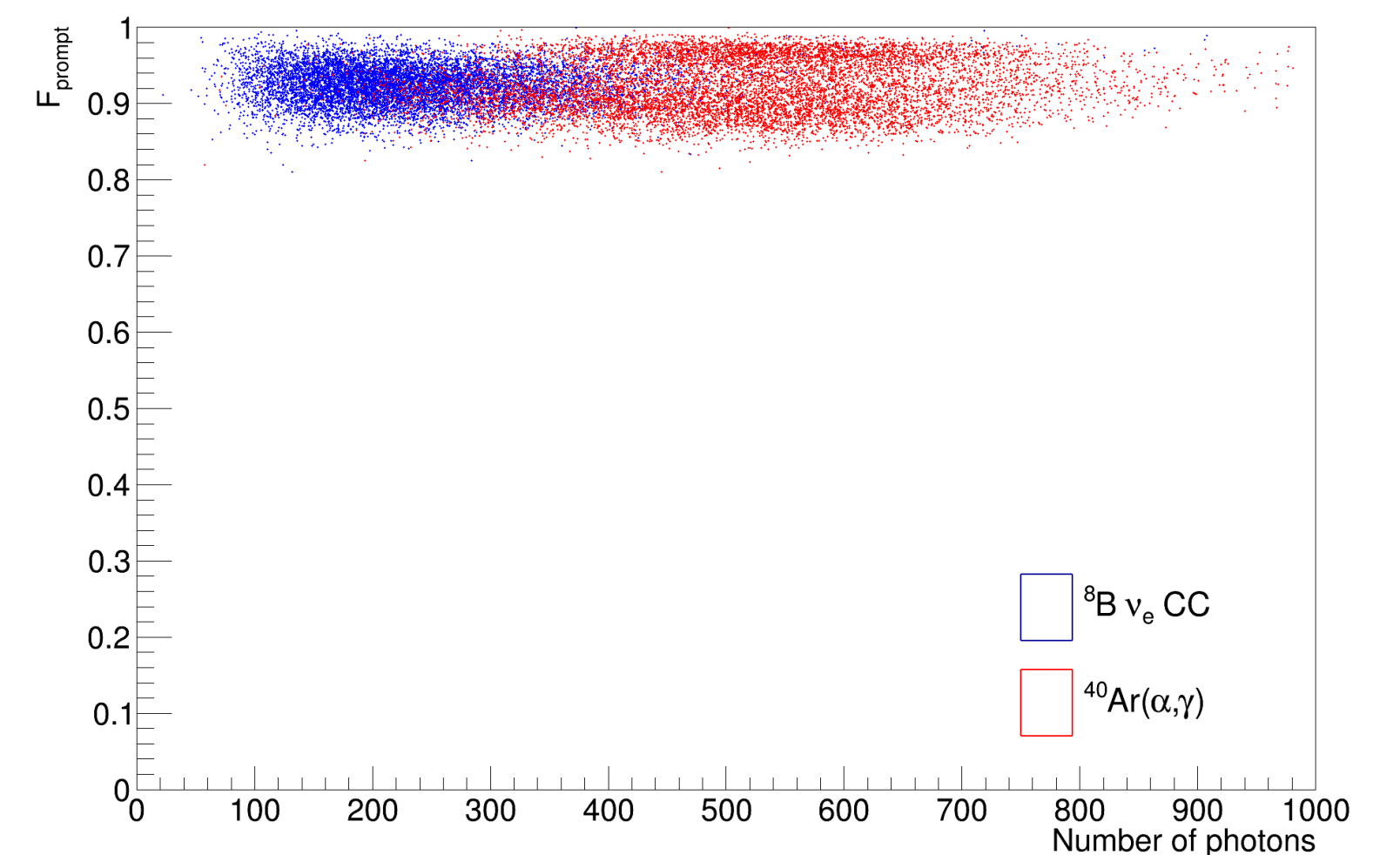
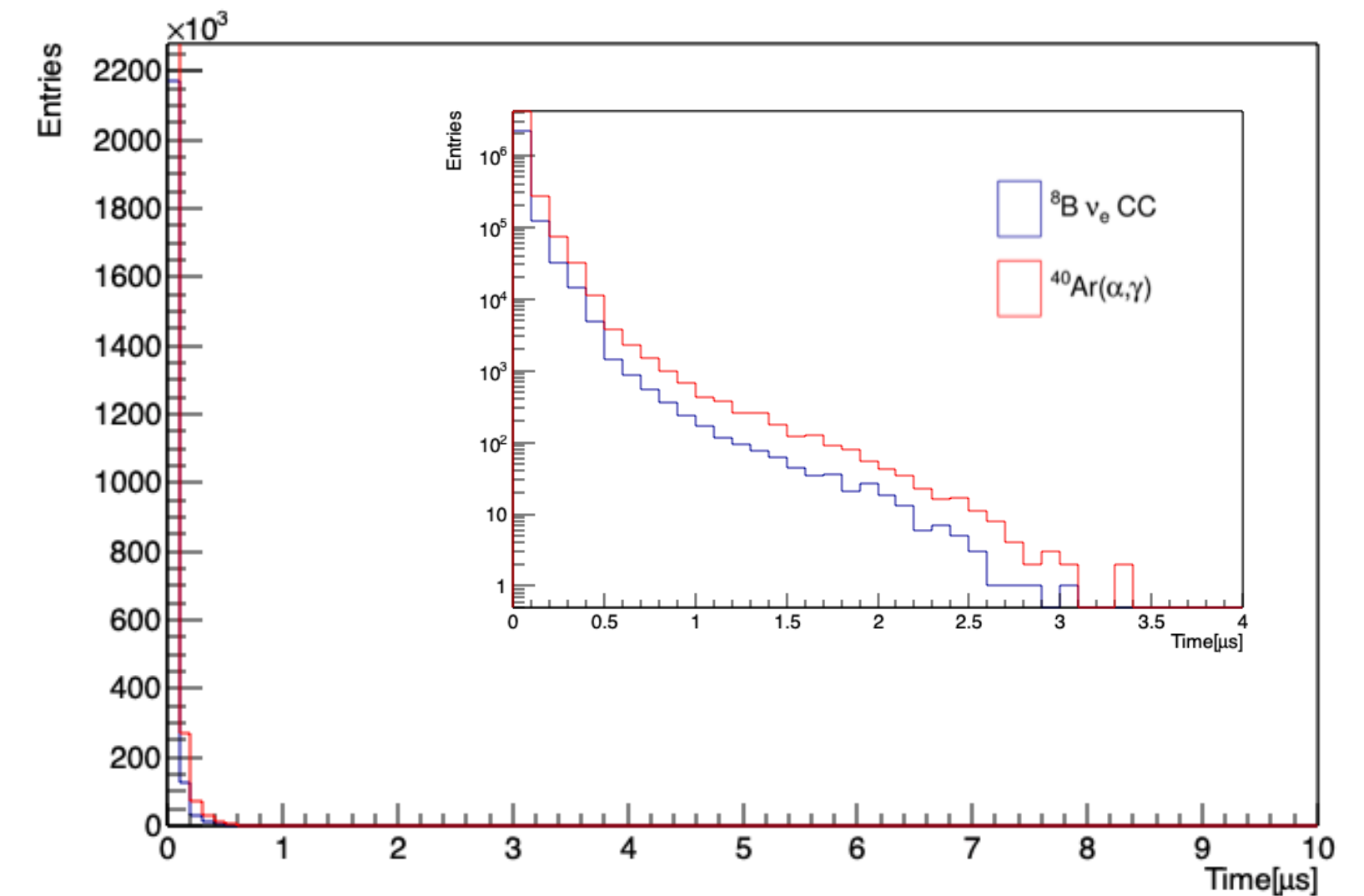
Adding xenon

- Doping with Xe changes scintillation timings and effectively removes the scintillation slow component

➔ f_{prompt} cut impossible

- Need to find alternative rejection method.

- Accumulation of $^{40}\text{Ar}(\alpha, \gamma)$ events near the cathode can be used to explore other discriminating observables.

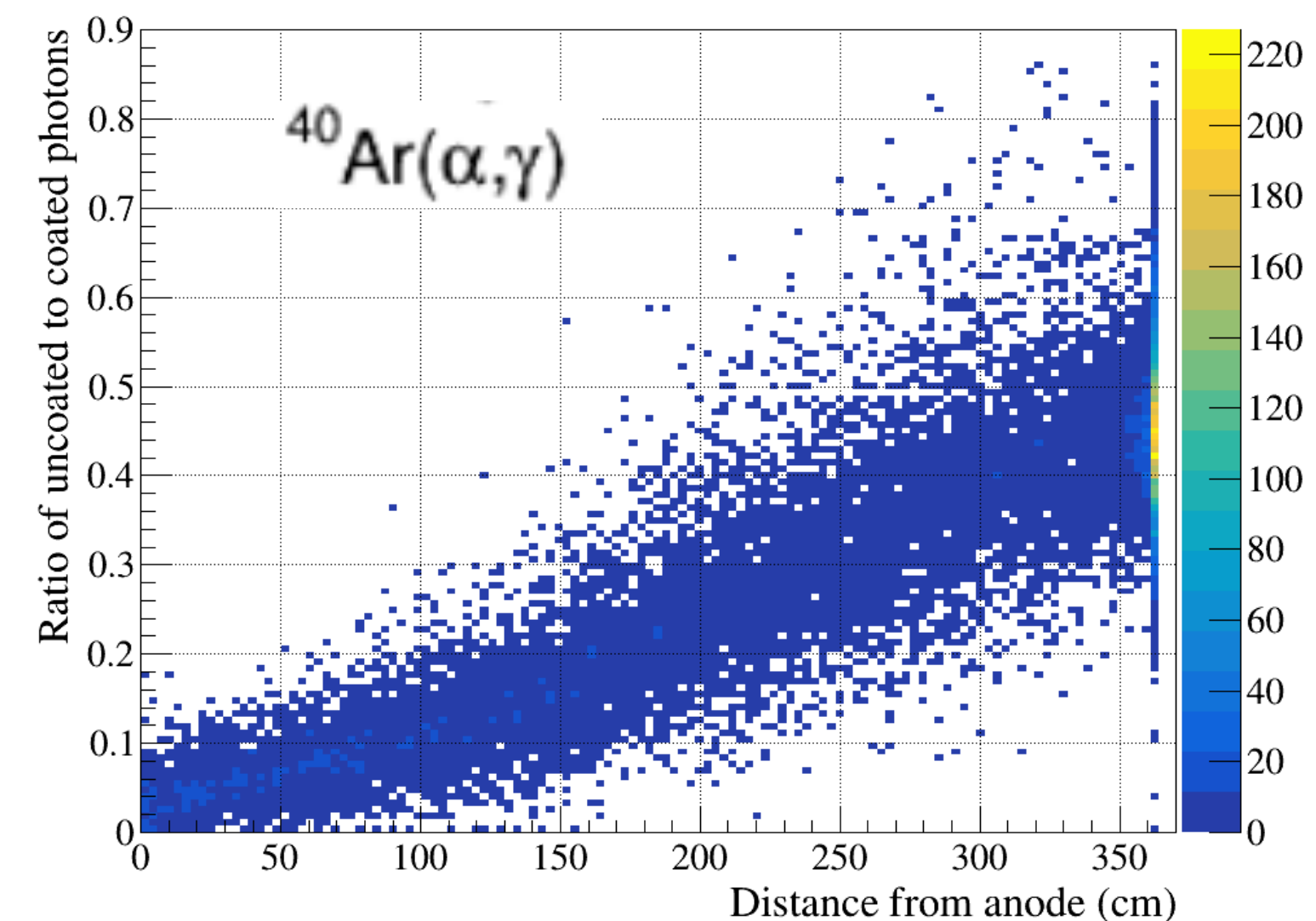
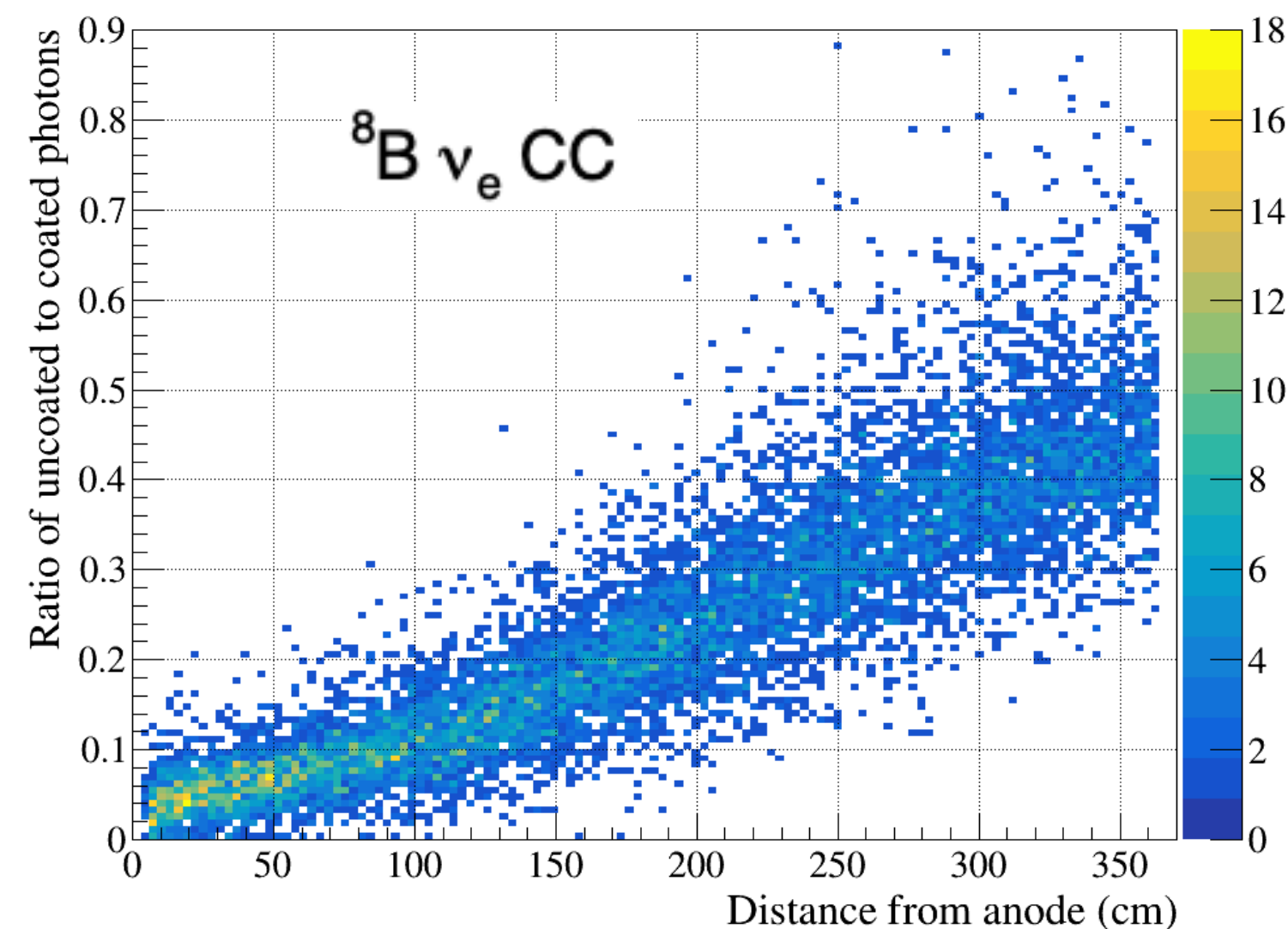
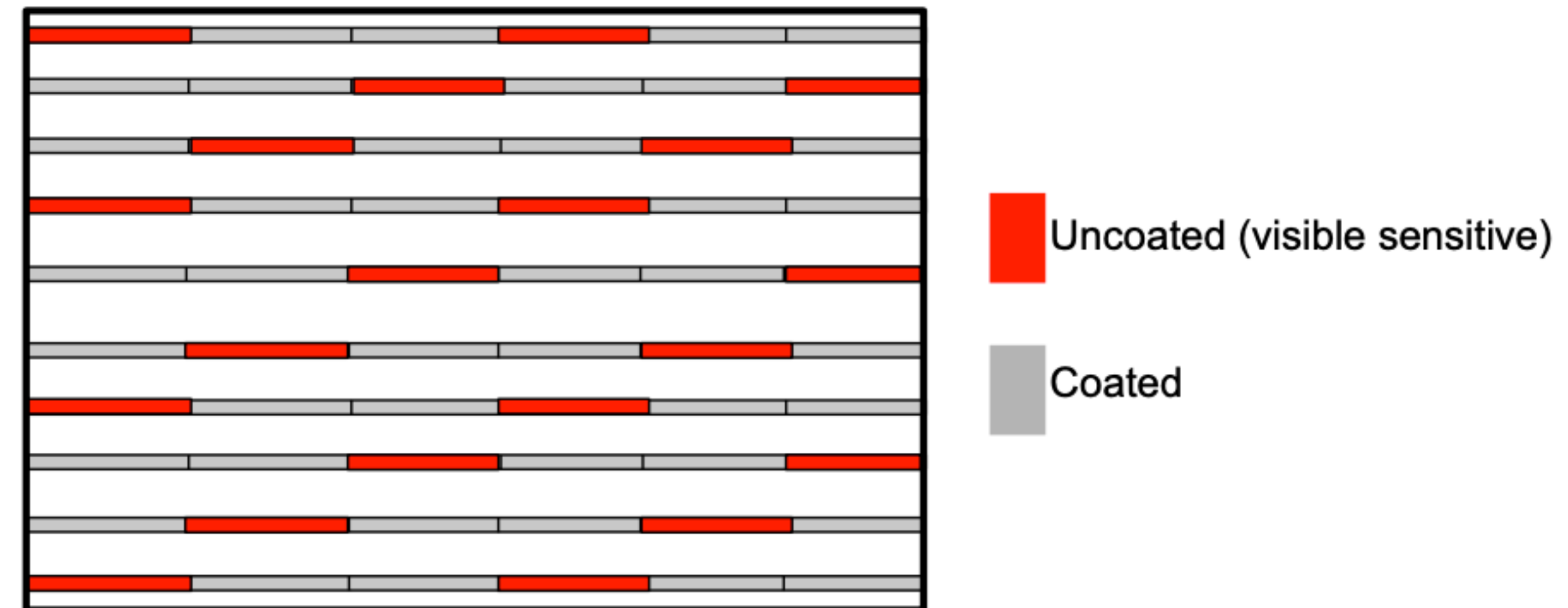


Exploring other methods

Other methods

Ratio of visible to VUV light

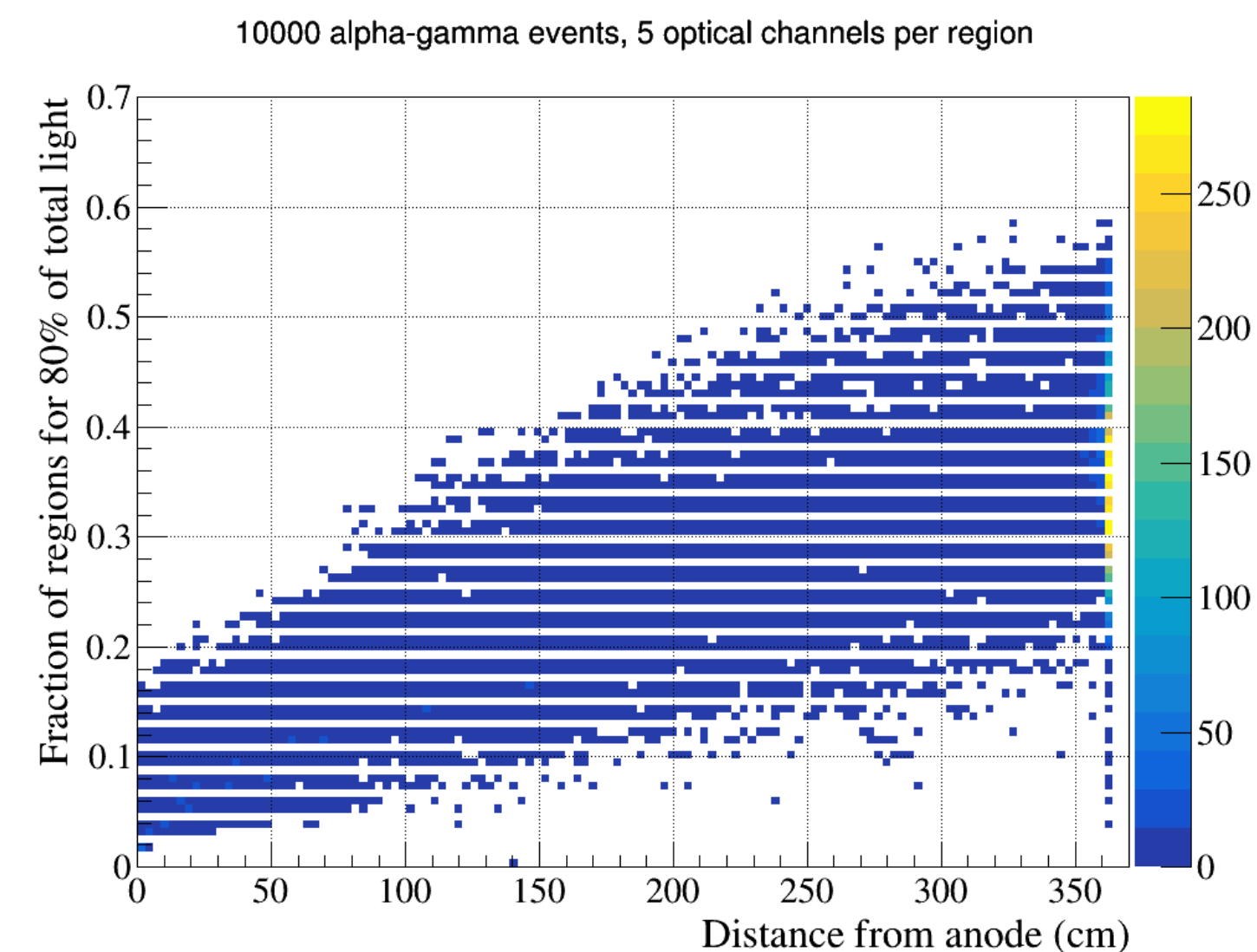
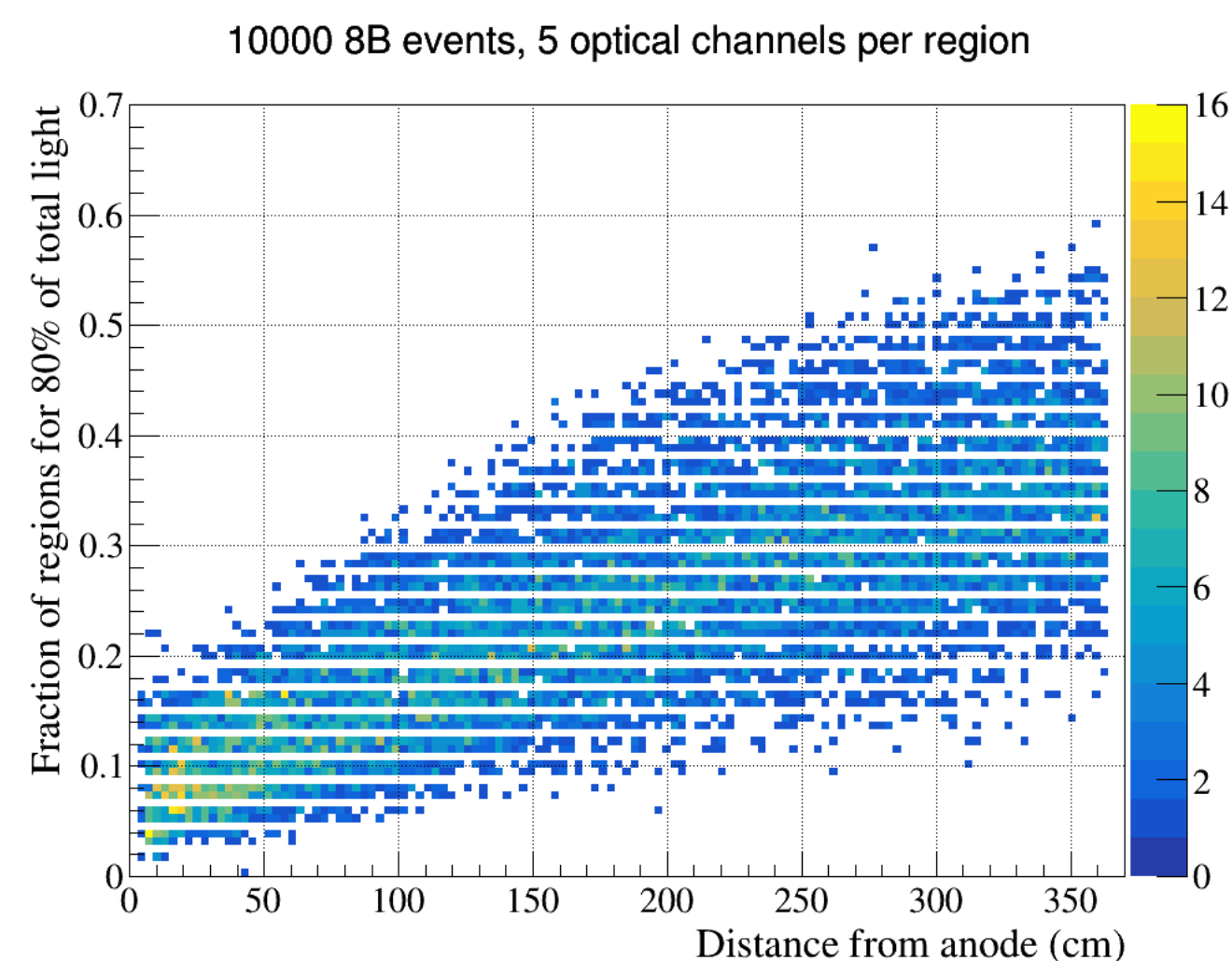
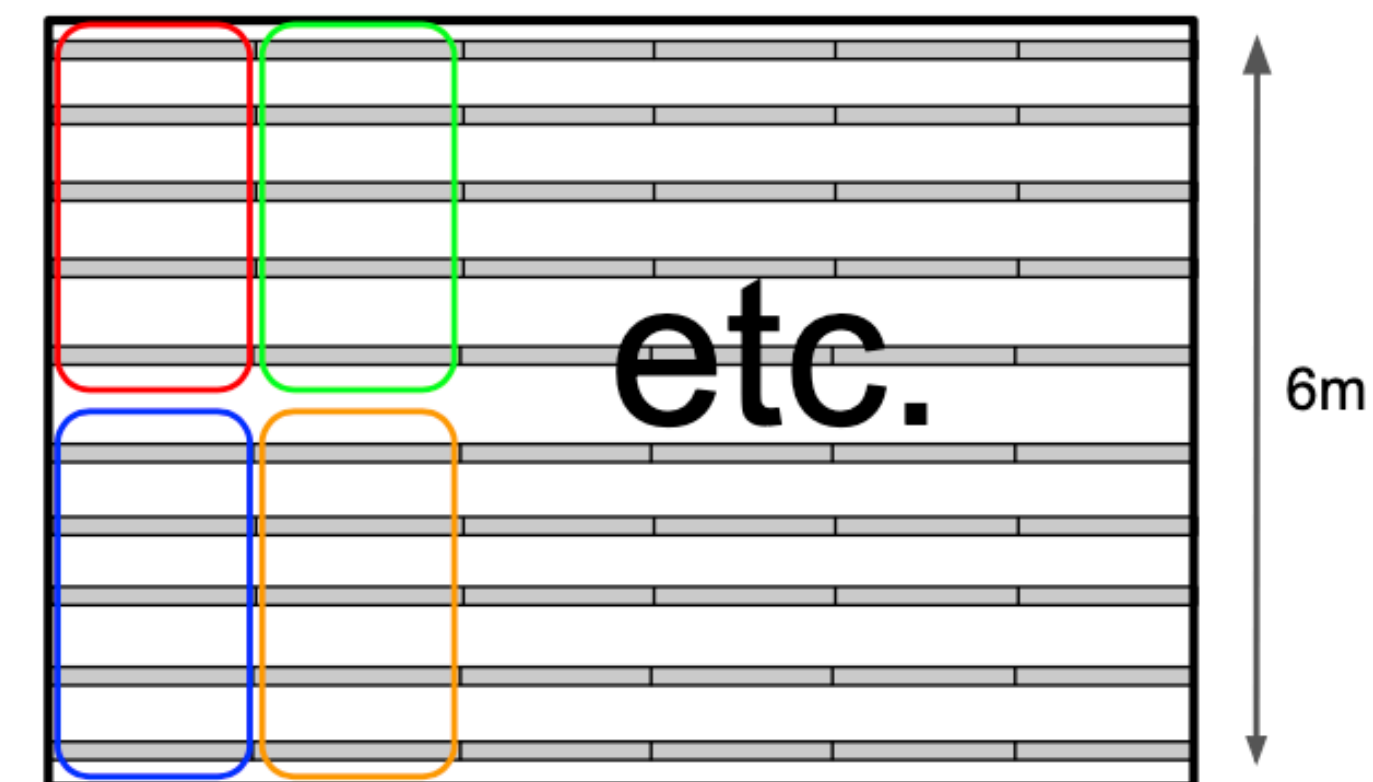
- Have some photon detectors not coated with wavelength shifter.
➡ Sensitive only to visible light reflected off of foils
- Coated detectors could be sensitive to both depending on coating
- Events close to CPA will have a higher fraction of visible light detected



Other methods

Light “spreading”

- Divide the optical detectors into small groups or “regions”
- Record the number of “regions” needed to detect 80% of the total light from an event
- Events close to the CPAs will have a more spread out light pattern
- Can vary number of channels per region to optimise results



Summary and next steps

- DUNE has great potential for solar neutrinos studies
- ^{222}Rn induced background can be challenging, but scintillation light is a promising way to mitigate it
- Many ideas to be explored:
 - Non-PSD rejection methods to account for Xe doping
 - Combined cuts
 - Tune selection for hep flux
 - Add vertical drift configuration



The University of Manchester

Thank you!

Questions?